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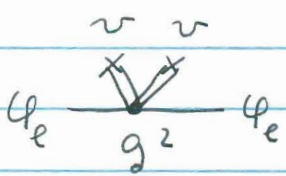
Back to:  $\phi_e, \phi_n$   $G = Z_2 \times Z_2$

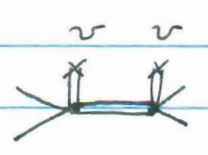
$\frac{m_n^2}{\lambda v^2} \gg m^2$   
 $\lambda \gg g^2$

$$\mathcal{L}_{UV}(\phi_e, \phi_n) = \frac{1}{2} (\partial\phi_e)^2 - \frac{1}{2} m^2 \phi_e^2 + \frac{1}{2} (\partial\phi_n)^2 - \frac{\lambda}{4} (\phi_n^2 - v^2)^2 - g^2 \phi_e^2 \phi_n^2 + \dots$$

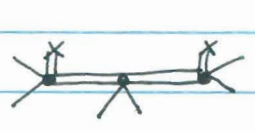
$$\mathcal{L}_{IR}(\phi_e) = \frac{1}{2} (\partial\phi_e)^2 - \frac{1}{2} (m^2 + 2g^2 v^2) \phi_e^2 - \frac{g^4}{\lambda} \phi_e^4 + \dots$$

Estimate  $\phi_e^6, \dots, \phi_e^{2n}$

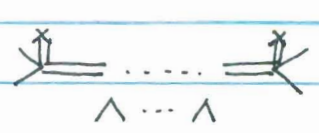

 $\approx g^2 v^2 \phi_e^2$


 $\approx \frac{(g^2)^2 v^2}{\lambda v^2} \phi_e^4$ 

$$\frac{g^4}{\lambda}$$


 $\approx \frac{(g^2)^3 v^2}{(\lambda v^2)^2} \phi_e^6$ 

$$\frac{m_n^2}{\lambda}$$


 $\approx \frac{g^{2n} v^2}{(\lambda v^2)^{n-1}} \phi_e^{2n}$ 

$$= \frac{g^{2n}}{\lambda} \frac{1}{(\lambda v^2)^{n-2}} \equiv \frac{g^{2n}}{\lambda} \frac{1}{(m_n^2)^{n-2}}$$

## Remarks

\*  $\mathcal{L}_{IR}(\varphi_e)$  can be diagrammatically obtained by summing all diagrams with heavy tree-level internal lines.

These are 1PI with respect to the light lines

\* The power counting is not completely fixed by dimensional analysis

$$m_h^2 \approx \lambda v^2$$

We could have expected, on dimensional grounds

$$\frac{\varphi^{2n}}{(m_h^2)^{n-2}}$$

Indeed we have:  $\frac{g^{2n}}{\lambda (m_h^2)^{n-2}} \varphi^{2n}$

The precise powers of  $g$  and  $\lambda$  can only come from the UV theory.

The EFT description breaks down at

$$E \approx m_h$$

since reaching this energy we start producing the heavy particles not included in  $\mathcal{L}_{IR}$ .

on the other hand, if  $g^2 \ll 1$  and  $g^2 \ll \Lambda$   
The effects in the irrelevant operators are  
suppressed by a larger effective scale

$$\Lambda_{\text{eff}}^2 \approx \left( \frac{\Lambda}{g^{2n}} \right)^{\frac{1}{n-2}} m_h^2 \gg m_h^2$$

This is like in the Fermi theory where  
the correct power counting gives

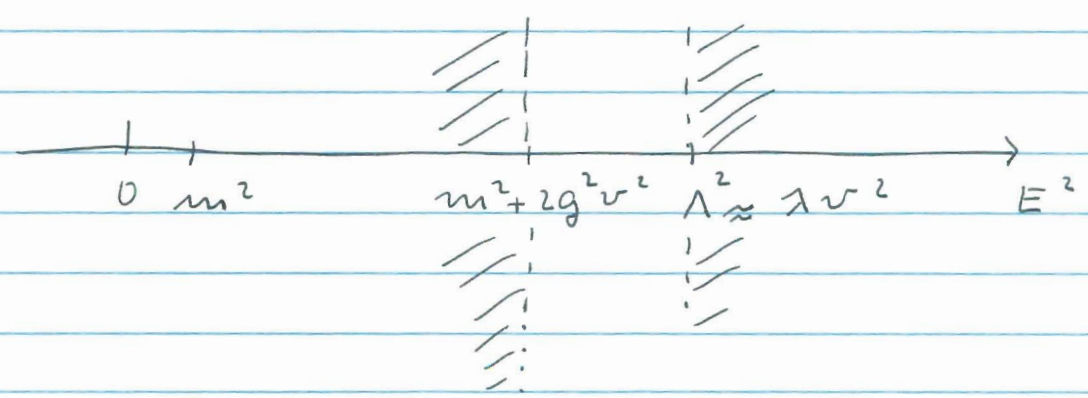
$$\frac{g^2}{M_W^2} \bar{\Psi} \Psi \bar{\Psi} \Psi \approx G_F \bar{\Psi} \Psi \bar{\Psi} \Psi$$

The IR description breaks at  $E \approx M_W$   
but  $\sqrt{1/G_F} \approx v_{EW} = 250 \text{ GeV} > M_{W,Z}$ .

\* A relevant operator has been generated

$$m^2 \rightarrow m^2 + 2g^2 v^2$$

Even if  $g^2 \ll 1$   $g^2 v^2 \gg m^2$  thus  
restricting the window where our EFT  
applies



Use of classical equations of motion is equivalent to evaluation of tree-level diagrams with heavy internal lines. Which is better / more practical?

→ I prefer e.o.m. since it automatically accounts for combinatorial factors.  
 Example: back to

$$\mathcal{L}(\varphi, \Psi) = \frac{1}{2} (\partial\varphi)^2 - \frac{1}{2} M^2 \varphi^2 + \bar{\Psi}(i\gamma - m)\Psi - g\varphi\bar{\Psi}\Psi$$

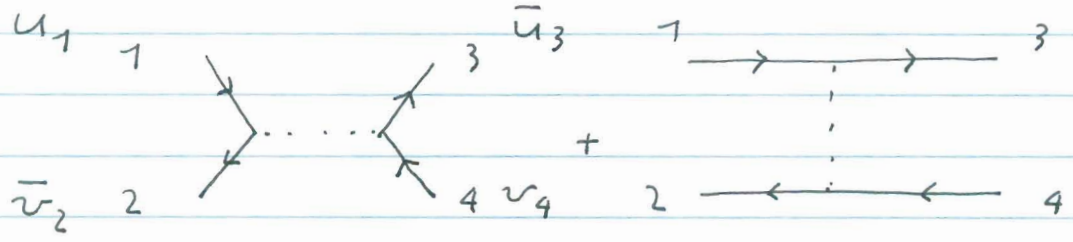
$$M^2 \gg m^2$$

Call the particles described by  $\Psi$  electrons  $e$  and consider

$$e^+e^- \rightarrow e^+e^-$$

to derive  $\mathcal{L}_{IR}(\Psi)$ .

At the tree level we have 2 diagrams



as usual  $s = (k_1 + k_2)^2$   $t = (k_1 - k_3)^2$



and

$$A(e^+e^- \rightarrow e^+e^-) =$$

$$(-ig^2) \left[ \frac{\bar{v}_2 u_1 \bar{u}_3 v_4}{s - M^2} + \frac{\bar{u}_3 u_1 \bar{v}_2 v_4}{t - M^2} \right]$$

$$\xrightarrow{s, t \ll M^2} + \frac{ig^2}{M^2} (\bar{v}_2 u_1 \bar{u}_3 v_4 + \bar{u}_3 u_1 \bar{v}_2 v_4)$$

Which Lagrangian reproduces this?

If we are sloppy, we write

$$\mathcal{L}_{IR} = \dots + \frac{g^2}{M^2} \bar{\Psi} \Psi \bar{\Psi} \Psi + \dots$$

missing a factor  $\frac{1}{2}$  here

Indeed:

$\langle 3 \bar{4} | \bar{\Psi} \Psi \bar{\Psi} \Psi | 1 \bar{2} \rangle$  produces 4 contributions:

2 s-channel:  $\langle 3 \bar{4} | \bar{\Psi} \Psi \bar{\Psi} \Psi | 1 \bar{2} \rangle$

$$\langle 3 \bar{4} | \bar{\Psi} \Psi \bar{\Psi} \Psi | 1 \bar{2} \rangle$$

2 t-channel:  $\langle 3 \bar{4} | \bar{\Psi} \Psi \bar{\Psi} \Psi | 1 \bar{2} \rangle$

$$\langle 3 \bar{4} | \bar{\Psi} \Psi \bar{\Psi} \Psi | 1 \bar{2} \rangle$$

and we should put a factor  $\frac{1}{2}$ .

This comes automatically by using the e.o.m.

## SUMMARY

- For a rapid estimate of power counting diagrams are O.K.
- To fix coefficients e.o.m. are better unless you are very good at computing diagrams

Exercise

Yukawa theory with  $\lambda \phi^4$  interaction:

$$\mathcal{L}_{UV} = \frac{1}{2} (\partial\phi)^2 - \frac{1}{2} M^2 \phi^2 - \frac{\lambda}{4} \phi^4 + \bar{\Psi} (i\cancel{\partial} - m) \Psi - g \phi \bar{\Psi} \Psi$$

Derive  $\mathcal{L}_{IR}$  at tree level keeping  $O(\lambda)$  terms  
 $m \ll M$ , neglecting the kinetic term  
EOM in the static limit:

$$M^2 \phi^2 + \lambda \phi^3 + g \bar{\Psi} \Psi = 0$$

0th order in  $\lambda$ :

$$\phi_0 = - \frac{g \bar{\Psi} \Psi}{M^2} = \frac{J}{M^2} \quad J \equiv -g \bar{\Psi} \Psi$$

1st order:

$$\phi_1 = \phi_0 + \lambda \delta\phi$$

$$M^2 \cancel{\phi_0} + \lambda M^2 \delta\phi + \lambda \phi_0^3 - \cancel{J} = 0$$

$$\delta\phi = - \frac{\phi_0^3}{M^2}$$

$$\phi_1 = \frac{J}{M^2} \left( 1 - \lambda \frac{J^2}{M^6} + \dots \right)$$

$$\delta\phi = - \frac{J^3}{M^8}$$

Back into  $\mathcal{L}_{UV}$ , neglecting the k. term

$$\mathcal{L}_{IR} = - \frac{1}{2} M^2 \frac{J^2}{M^4} \left( 1 - 2 \lambda \frac{J^2}{M^6} + \dots \right) - \frac{\lambda}{4} \frac{J^4}{M^8} + \bar{\Psi} (i\cancel{\partial} - m) \Psi - \frac{g \bar{\Psi} \Psi}{+J} \cdot \frac{J}{M^2} \left( 1 - \lambda \frac{J^2}{M^6} + \dots \right)$$

$$= \frac{J^2}{M^2} \left( -\frac{1}{2} + 1 \right) + \lambda \frac{J^4}{M^8} \left( +1 - \frac{1}{4} - 1 \right) + \dots$$
$$= \frac{1}{2} \frac{J^2}{M^2} - \frac{\lambda}{4} \frac{J^4}{M^8} + \dots = \frac{1}{2} \frac{g^2 (\bar{\Psi} \Psi)^2}{M^2} - \frac{\lambda g^4 (\bar{\Psi} \Psi)^4}{4 M^8} + \dots$$

~~$g^6 (\bar{\Psi} \Psi)^6 \lambda^2$~~   $\lambda^{n-1} g^{2n} \frac{(\bar{\Psi} \Psi)^{2n}}{M^{2(3n-2)}}$

and/or

but how we proceed if  $\lambda/g \gg 1$ ?

We cannot rely on perturbation theory

1. Assume that at  $E \ll M$  the relevant particles are still described by  $\Psi$   
 Not guaranteed. In a strong coupling regime particles other than  $\psi$  and  $\bar{\psi}$  can be formed. Some of them can be light. For instance in QCD the fundamental fields are quarks  $q$  and gluons  $G$  but in the low-energy regime we have light particles such as  $\pi$ 's, etc...

2 Assume Lorentz invariance

$$\mathcal{L}_{IR} = \bar{\Psi}(i\gamma^\mu \partial_\mu - m')\Psi + \frac{C_4^{(1)}}{\Lambda^2} (\bar{\Psi}\Psi)^2 + \dots$$

$$+ \frac{C_4^{(2)}}{\Lambda^2} (\bar{\Psi}\gamma^\mu\Psi)(\bar{\Psi}\gamma_\mu\Psi) + \dots$$

$$- \frac{C_4^{(3)}}{\Lambda^2} (\bar{\Psi}\sigma^{\mu\nu}\Psi)(\bar{\Psi}\sigma_{\mu\nu}\Psi) + \dots$$



Exercise

Spontaneously broken U(1) theory

$\phi$  complex scalar  $Q(\phi) = +1$

$\psi$  Dirac fermion  $Q(\psi) = +1$

$$\mathcal{L}_{UV} = \bar{\psi}(i\not{\partial} - m)\psi + (D_\mu\phi)^\dagger D^\mu\phi - \frac{1}{4} F_{\mu\nu} F^{\mu\nu} - \frac{\lambda}{4} (|\phi|^2 - v^2)^2$$

$$D_\mu\phi = (\partial_\mu + ig A_\mu)\phi$$

$$D_\mu\psi = (\partial_\mu + ig A_\mu)\psi$$

Determine  $\mathcal{L}_{IR}$  at tree-level in the static approximation in the regimes

- ①  $g^2 v^2 \gg m^2, \lambda v^2$
- ②  $\lambda v^2 \gg m^2, g^2 v^2$

Mass spectrum

fermion  $\psi$  has mass  $m$

$\lambda > 0$   $\langle \phi \rangle = \frac{v}{\sqrt{2}}$  : gauge symmetry is spontaneously broken

Use new fields:

$$\phi = \frac{\rho}{\sqrt{2}} e^{i \frac{\xi}{v}}$$

$$V = \frac{\lambda}{16} (\rho^2 - v^2)^2 \quad \xi - \text{independent}$$

$$V_{\xi} = 0 \quad V_{\rho} = \frac{\lambda}{4} (\rho^2 - v^2) \rho = 0$$

$\rho = 0$        $\rho = v$   
local      local minimum  
max

$$V_{\rho\rho} = \frac{\lambda}{4} (\rho^2 - v^2) + \frac{\lambda}{2} \rho^2$$

$$V_{\rho\rho}|_0 = -\frac{\lambda}{4} v^2 < 0 \quad V_{\rho\rho}|_v = \frac{\lambda}{2} v^2 > 0$$

$$m_{\xi} = 0 \quad \text{redefine } \rho = h + v$$

$$m_h^2 = \frac{\lambda}{2} v^2$$

Finally

$$(D_{\mu} \phi)^{\dagger} D^{\mu} \phi = \left| (\partial_{\mu} + i g A_{\mu}) \frac{h+v}{\sqrt{2}} e^{i \xi/v} \right|^2$$

$\rightarrow$  unitary gauge  $\xi = 0$

$$\left| \partial_{\mu} \frac{h}{\sqrt{2}} + i g A_{\mu} \left( \frac{h+v}{\sqrt{2}} \right) \right|^2 =$$

$$= \frac{1}{2} \partial_{\mu} h \partial^{\mu} h + \frac{g^2}{2} (h+v)^2 A_{\mu} A^{\mu} \rightarrow m_A^2 = g^2 v^2$$

SUMMARY

field	$\psi$	$\xi$	$h$	$A$
mass <sup>2</sup>	$m^2$	eaten up by $A_\mu$	$\frac{1}{2} \lambda v^2$	$g^2 v^2$

$$\mathcal{L}_{UV} = -\frac{1}{4} F_{\mu\nu} F^{\mu\nu} + \frac{1}{2} \partial_\mu h \partial^\mu h + \bar{\psi} i \gamma^\mu \partial_\mu \psi - \bar{\psi} m \psi$$

$$- g A_\mu \bar{\psi} \gamma^\mu \psi + \frac{g^2}{2} (h+v)^2 A_\mu A^\mu$$

$$- \frac{\lambda}{16} ((h+v)^2 - v^2)^2$$

①  $g^2 v^2 \gg m^2, \lambda v^2$

$\rightarrow m_A$  is the heaviest mass

e.o.m. for  $A_\mu$  in the static limit (i.e. neglecting ~~the~~  $F_{\mu\nu} F^{\mu\nu}$ )

$$- g \bar{\psi} \gamma^\mu \psi + g^2 (h+v)^2 A^\mu = 0$$

$$A_\mu = \frac{g \bar{\psi} \gamma^\mu \psi}{g^2 (h+v)^2} = \frac{g}{m_A^2} \cdot \frac{\bar{\psi} \gamma^\mu \psi}{(1 + \frac{h}{v})^2}$$

back into the Lagrangian  $\mathcal{L}_{UV}$ :

$$\mathcal{L}_{IR} = \frac{1}{2} \partial_\mu h \partial^\mu h + \bar{\psi} i \gamma^\mu \partial_\mu \psi - \bar{\psi} m \psi - V(h)$$

$$- \frac{g^2}{m_A^2} \frac{(\bar{\psi} \gamma^\mu \psi)^2}{(1 + \frac{h}{v})^2} + \frac{g^2}{2} \frac{(h+v)^2}{m_A^4} \frac{g^2 (\bar{\psi} \gamma^\mu \psi)^2}{(1 + \frac{h}{v})^4}$$



$$\mathcal{L}_{IR} = \frac{1}{2} \partial_\mu h \partial^\mu h + \bar{\Psi} i \gamma^\mu \partial_\mu \Psi - \bar{\Psi} m \Psi - V(h) - \frac{1}{2} \frac{g^2}{m_A^2} \frac{(\bar{\Psi} \gamma^\mu \Psi)^2}{\left(1 + \frac{h}{v}\right)^2}$$

how should we interpret the last term

$$\frac{h}{v} = \frac{g h}{g v} = \frac{g h}{m_A}$$

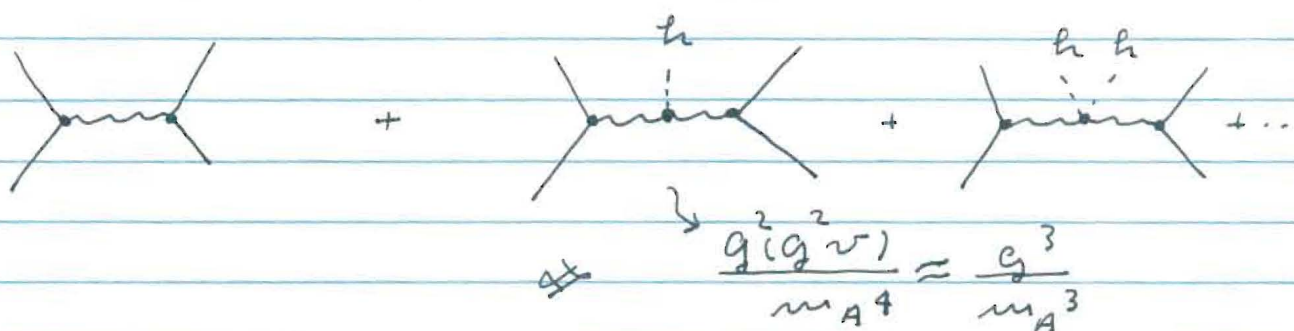
and we can expand in  $\frac{1}{m_A}$

$$\left(1 + \frac{h}{v}\right)^{-1} = 1 - \frac{g h}{m_A} + O(h^2)$$

The last term reads

$$-\frac{1}{2} \frac{g^2}{m_A^2} (\bar{\Psi} \gamma^\mu \Psi)^2 \left[ 1 - 2 \frac{g h}{m_A} + \dots \right]$$

In terms of diagrams



$$(2) \quad \lambda v^2 \gg g^2 v^2, m^2$$

$h$  is the heaviest state

$$V = \frac{\lambda}{16} (h^2 + 2v h)^2$$

$$= \frac{\lambda}{16} (h^4 + 4v h^3 + 4v^2 h^2)$$



$$g^2 A_\mu A^\mu - \frac{\lambda}{4} (\rho^2 - v^2) = 0$$

$$\rho = v \left( 1 + \frac{g^2 A^2}{m_h^2} \right)$$

$$g^2 A^2 - \frac{\lambda}{4} \left( v^2 \left( 1 + \frac{2g^2 A^2}{m_h^2} \right) - v^2 \right) = 0$$

better to use the variables:  $\rho \equiv h + v$   
 e.o.m. in the static limit

$$g^2 \rho A_\mu A^\mu - \frac{\lambda \rho}{4} (\rho^2 - v^2) = 0$$

we are working in the regime  $\lambda v^2 \gg g^2 v^2$   
 $\lambda \gg g^2$

we can solve this by expanding in series of  $g^2$   
 0th order

$$\rho^2 - v^2 = 0 \quad \rho = v \quad \leftrightarrow \quad h = 0$$

back into  $\mathcal{L}_{UV}$ :

$$\mathcal{L}_{IR} = -\frac{1}{4} F_{\mu\nu} F^{\mu\nu} + \bar{\Psi} i \gamma^\mu \partial_\mu \Psi - \bar{\Psi} m \Psi \\ - g A_\mu \bar{\Psi} \gamma^\mu \Psi + \frac{1}{2} m_H^2 A_\mu A^\mu$$

1st order

$$\rho = v (1 + x g^2)$$

$$g^2 v A_\mu A^\mu - \frac{\lambda v}{4} (v^2 (1 + 2x g^2) - v^2) = 0$$

$$v A_\mu A^\mu - \frac{\lambda v^3}{2} x = 0$$

$$x = \frac{A_\mu A^\mu}{m_H^2}$$

$$\rho = v \left( 1 + \frac{g^2}{m_H^2} A_\mu A^\mu + \dots \right)$$

$$\mathcal{L}_{IR} = \text{1st line} - g A_\mu \bar{\Psi} \gamma^\mu \Psi$$

$$+ \frac{g^2}{2} v^2 \left( 1 + \frac{2g^2}{m_H^2} A_\mu A^\mu + \dots \right) A_\nu A^\nu$$

$$- \frac{\lambda}{16} \left[ v^2 \left( 1 + \frac{2g^2}{m_H^2} A_\mu A^\mu + \dots \right) - v^2 \right]^2$$

$$\mathcal{L}_{IR} = 1^{st} \text{ line} - g A_\mu \bar{\Psi} \gamma^\mu \Psi + \frac{1}{2} m_A^2 A_\nu A^\nu$$

$$+ (A_\mu A^\mu)^2 \left[ \frac{g^4 v^2}{m_h^2} - \frac{1}{16} \frac{v^4 g^4}{m_h^4} \right]$$

$$\downarrow$$

$$\frac{1}{4} \frac{2 v^2}{16 m_h^4} \cdot 4 g^4 v^2$$

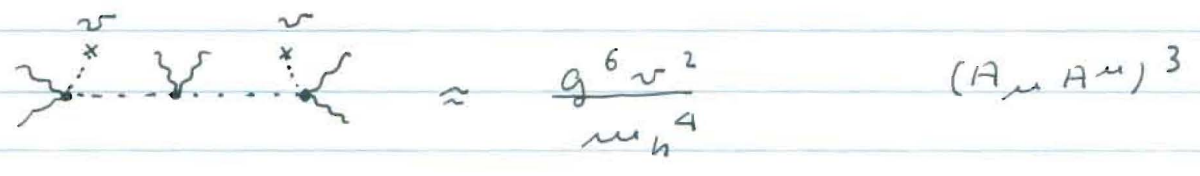
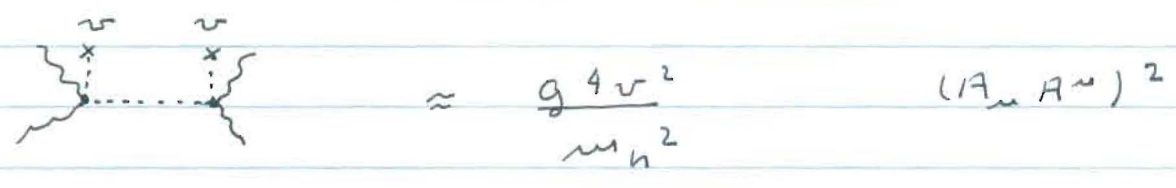
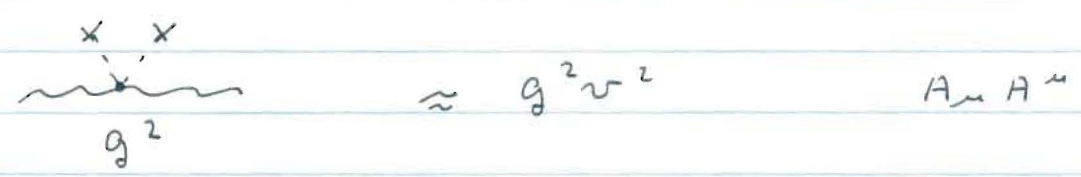
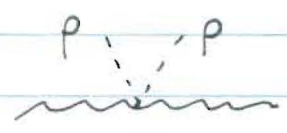
$$\equiv \frac{1}{2 m_h^2} g^4 v^2$$

$$\mathcal{L}_{IR} = 1^{st} \text{ line} - g A_\mu \bar{\Psi} \gamma^\mu \Psi + \frac{1}{2} m_A^2 A_\mu A^\mu$$

$$+ \frac{1}{2} \frac{g^4 v^2}{m_h^2} (A_\mu A^\mu)^2 + \dots$$

↑  
marginal operator

in terms of diagrams



Notice the peculiar power counting

$$\frac{g^{2n} v^2}{(m_h^2)^{n-1}} \approx (A_\mu A^\mu)^n$$



$$\text{CASE } g^2 v^2 \approx \lambda v^2 \gg m^2$$

$$\begin{cases} \bar{\Psi} \gamma^\mu \Psi + g \rho^2 A^\mu = 0 \\ g^2 A_\mu A^\mu - \frac{\lambda}{4} (\rho^2 - v^2) = 0 \end{cases}$$

1<sup>st</sup> gives  $A_\mu = \frac{\bar{\Psi} \gamma_\mu \Psi}{g \rho^2}$

in the 2<sup>nd</sup> gives:

$$\frac{g^2}{g^2 \rho^4} (\bar{\Psi} \gamma_\mu \Psi)^2 - \frac{\lambda}{4} (\rho^2 - v^2) = 0$$

$$(\bar{\Psi} \gamma_\mu \Psi)^2 = \frac{\lambda}{4} \rho^4 (\rho^2 - v^2)$$

To formally solve this, introduce a large expansion parameter  $u$

$$u (\bar{\Psi} \gamma_\mu \Psi)^2 \equiv u J^2 = \frac{\lambda}{4} \rho^4 (\rho^2 - v^2)$$

Expanding the solution in powers of  $u$ :

$$\rho = v (1 + x u)$$

$$J^2 = \frac{\lambda}{4} v^6 (1 + 4x u)^2 x^2 v^6 = \frac{\lambda}{2} x^2 v^6$$

$$\rightarrow x = \frac{2}{\lambda} \frac{J^2}{v^6}$$

$$\rho \equiv v \left( 1 + \frac{2}{\lambda} \frac{J^2}{v^6} + \dots \right) \quad A_\mu = \frac{J_\mu}{g v^2 \left( 1 + \frac{4}{\lambda} \frac{J^2}{v^6} \right)}$$

$$A_\mu \equiv \frac{J_\mu}{g v^2} \left( 1 - \frac{4}{\lambda} \frac{J^2}{v^6} + \dots \right)$$



$$\mathcal{L}_{IR} = -\frac{1}{4} F_{\mu\nu} F^{\mu\nu} + \bar{\Psi} i \gamma^\mu \partial_\mu \Psi - \bar{\Psi} m \Psi$$

$$- \cancel{g} \frac{J^2}{g v^2} \left( 1 - \frac{4}{\lambda} \frac{J^2}{v^6} + \dots \right)$$

$$+ \frac{g^2 \cancel{v^2}}{2} \left( 1 + \frac{4}{\lambda} \frac{J^2}{v^6} + \dots \right) \frac{J^2}{\cancel{g^2 v^4} 2}$$

$$- \frac{\lambda}{16} \left[ v^2 \left( \cancel{1} + \frac{4}{\lambda} \frac{J^2}{v^6} + \dots \right) - \cancel{v^2} \right]^2$$

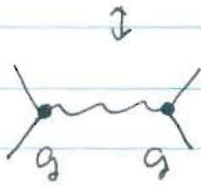
= 1<sup>st</sup> line

$$+ \frac{J^2}{v^2} \left( -1 + \frac{1}{2} \right)$$

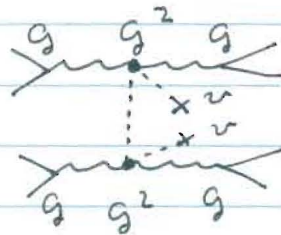
$$+ \frac{J^4}{\lambda v^8} \left( +4 + 2 - \frac{16}{16} \right)$$

= 1<sup>st</sup> line

$$- \frac{1}{2v^2} J^2 + \frac{5}{\lambda v^8} J^4 + \dots$$



$$\frac{g^2}{g^2 v^2} \approx \frac{1}{v^2}$$



$$\approx \frac{g^8 v^2}{m_H^8 m_H^2}$$

$$\approx \frac{g^8 \cancel{v^2}}{g^8 v^8 \lambda v^2}$$

$$\approx \frac{1}{\lambda v^8}$$

## Derivation of Fermi Lagrangian

Recap of the SM

① Gauge symmetry  $SU(3) \times SU(2) \times U(1)$

② Field content

- gauge vector bosons  $W_\mu^a, B_\mu, G_\mu$

- scalars:  $\phi \sim (1, 2, +1/2)$

- fermions: 3 generations of quarks and leptons

$$q_L \sim (3, 2, +1/6) = \begin{bmatrix} u_L \\ d_L \end{bmatrix}$$

$$u_R \sim (3, 1, +2/3)$$

$$d_R \sim (3, 1, -1/3)$$

$$l_L \sim (1, 2, -1/2) = \begin{bmatrix} \nu_L \\ e_L \end{bmatrix}$$

$$e_R \sim (1, 1, -1)$$

③ Renormalizability (only relevant and marginal op.)

$$\textcircled{1} + \textcircled{2} + \textcircled{3} \rightarrow \mathcal{L}_{SM} =$$

$$= -\frac{1}{4} F^2 + i \bar{\Psi} \gamma^\mu D_\mu \Psi$$

$$+ (D_\mu \phi)^\dagger D^\mu \phi - V(\phi^\dagger \phi)$$

$$- (\bar{\Psi} \psi \phi \Psi + \text{h.c.})$$

More in detail:

$$i \bar{\Psi} \gamma^\mu D_\mu \Psi = \text{kinetic terms} - e A_\mu J^\mu_{em}$$

$$- \frac{g}{2} (W_\mu^+ J^{-\mu} + W_\mu^- J^{+\mu})$$

$$- \sqrt{g^2 + g'^2} Z_\mu J^\mu_Z$$

+ gluon interactions

$$- g_s G_\mu^a \sum_i \bar{q} \gamma^\mu t^a q$$

where  $J^{\mu em} = \sum_f \bar{f} \gamma^\mu Q_f f$   
 $= -\bar{e} \gamma^\mu e + \frac{2}{3} \bar{u} \gamma^\mu u - \frac{1}{3} \bar{d} \gamma^\mu d$

$J^{+\mu} = \bar{d}_L \gamma^\mu u_L + \bar{e}_L \gamma^\mu \nu_L$   
 $J^{-\mu} = \bar{u}_L \gamma^\mu d_L + \bar{\nu}_L \gamma^\mu e_L$

$J_{3L}^{\mu} = \sum_f \bar{f} T_{3L}^f f$   
 $= \frac{1}{2} (\bar{u}_L \gamma^\mu u_L - \bar{d}_L \gamma^\mu d_L)$   
 $+ \frac{1}{2} (\bar{\nu}_L \gamma^\mu \nu_L - \bar{e}_L \gamma^\mu e_L)$

$J_2^{\mu} = J_{3L}^{\mu} - S_{\theta}^2 J^{\mu em}$

$\tan \theta \equiv g'/g$

$m_A = m_g = 0$        $m_W^2 = \frac{g^2 v^2}{4}$        $m_Z^2 = \frac{(g^2 + g'^2) v^2}{4}$

$\frac{m_W^2}{m_Z^2 \cos^2 \theta} = 1$

✳

Moreover, from  $(D_\mu \phi)^\dagger D^\mu \phi$  we get

$(D_\mu \phi)^\dagger D^\mu \phi = m_W^2 W_\mu^+ W^{-\mu} + \frac{1}{2} m_Z^2 Z_\mu Z^\mu$



The Fermi Lagrangian refers to an IR EFT derived in the limit  $E \ll m_{W,Z}$

1 at the TL

2 in the static approximation

3 at LO in  $g$  expansion

$$\text{Since } F_{\mu\nu} F^{\mu\nu} \sim \underbrace{(\partial A)^2}_{\text{neglected (2)}} + \underbrace{g A^2 \partial A}_{\text{neglected (2)}} + \underbrace{g^2 A^4}_{\text{neglected (3)}}$$

Relevant terms:

$$\begin{aligned} \mathcal{L}_{SM} = & \frac{1}{2} m_W^2 W_\mu^+ W^{-\mu} + \frac{1}{2} m_Z^2 Z_\mu Z^\mu \\ & - \frac{g}{2} (W_\mu^+ J^{-\mu} + W_\mu^- J^{+\mu}) \\ & - \sqrt{g^2 + g'^2} Z_\mu J_Z^\mu + \dots \end{aligned}$$

e.o.m.

$$W_\mu^\pm \quad m_W^2 W^\mp{}^\mu - \frac{g}{2} J^\mp{}^\mu = 0$$

$$Z_\mu \quad m_Z^2 Z^\mu - \sqrt{g^2 + g'^2} J_Z^\mu = 0$$

$$W^\mp{}^\mu = \frac{g}{2 m_W^2} J^\mp{}^\mu$$

$$Z^\mu = \frac{\sqrt{g^2 + g'^2}}{m_Z^2} J_Z^\mu$$

back into  $\mathcal{L}_{SM}$ :

$$\begin{aligned} \mathcal{L}_{Fermi} = & \frac{m_W^2 g^2}{2 m_W^4} J^+ J^- + \frac{1}{2} \frac{m_Z^2 (g^2 + g'^2)}{m_Z^4} J_Z^2 \\ & - \frac{g}{2} \frac{g}{2 m_W^2} 2 J^+ J^- - \frac{(g^2 + g'^2)}{m_Z^2} J_Z^2 \end{aligned}$$



original Fermi Lagrangian was of  $V \times V$  type

1957 Sudarshan Marshak

$(V-A) \times (V-A)$

Feynman Gell-Mann

$$= -\frac{g^2}{2m_W^2} J^+ J^- - \frac{(g^2 + g'^2)}{2m_Z^2} J_Z^2 + \dots = \mathcal{L}_{\text{Fermi}}$$

$$\frac{g^2}{2m_W^2} = \frac{4g^2}{2g^2 v^2} = \frac{2}{v^2}$$

$$\frac{(g^2 + g'^2)}{2m_Z^2} = \frac{4(g^2 + g'^2)}{2(g^2 + g'^2)v^2} = \frac{2}{v^2}$$

Define  $\frac{G_F}{\sqrt{2}} = \frac{g^2}{8m_W^2} = \frac{2}{v^2} = 2\sqrt{2} G_F$

$$\mathcal{L}_{\text{Fermi}} = -2\sqrt{2} G_F (J^+ J^- + J_Z^2)$$

For instance, in the lepton sector

$$J^+ = \bar{e}_L \gamma_\mu \nu_{eL} + \dots = \frac{1}{2} \bar{e} \gamma_\mu (1 - \gamma_5) \nu + \dots$$

$$J^- = \bar{\nu}_{\mu L} \gamma_\mu \mu_L + \dots = \frac{1}{2} \bar{\nu}_\mu \gamma_\mu (1 - \gamma_5) \mu + \dots$$

$$\mathcal{L}_{\text{Fermi}} = -\frac{G_F}{\sqrt{2}} \bar{\nu}_\mu \gamma_\mu (1 - \gamma_5) \mu * \bar{e} \gamma_\mu (1 - \gamma_5) \nu + \dots$$

describes  $\mu$ -decay

### Remarks

→ Same weight between  $J^+ J^-$  and  $J_Z^2$   
as a consequence of  $\frac{m_W^2}{m_Z^2 \cos^2 \theta} = 1$

↔ EW symmetry broken by  $\phi \sim (2, +1/2)$